

Welcome  
to 8.033!



## Relation to other courses

8.20 covers 60% of material - either stay in 8.033 and coast through initially or take 8.224 (“Exploring Black Holes”). Taking *both* 8.20 and 8.224 waives the 8.033 requirement.

## Surgeon General’s Warning

- **False sense of security:** Many of you will may think you know it all from books, other courses, *etc.* Don’t get lulled into false sense of security!
- **Unusual style:** Not everything is proven from the beginning (cf. 8.01 & 8.02). For SR, we’ll derive everything from Einstein’s 2 postulates, but for cosmo and GR we jump right into the middle and work out cool consequences from equations you’ll need to take at faith
- **Non-intuitive:** This may be your first real departure from intuitive physics.
- **Collaboration:** OK to collaborate on problem sets, but you *must* hand in your own work. Copying ==> trouble.
- **Night owls:** Like to stay up all night? View Friday as part of the weekend? Consider grad school & postdoc! No credit for late problem sets. Attendance crucial: saves you time, and beyond books, there’ll be  $\sim 30\%$  value-added that you’re responsible for on exams. If you miss a lecture, see friend or TA.
- **Unfamiliar math:** Calculus of variations
- **Lots of PowerPoint:** Keep me from going to fast and erasing too soon

How to do well in the course:

- Come to class
- Use study guide
- Do all psets
- Study with someone else
- Don't waste time on pset all-nighters with no help

I've never failed a student who really tried hard.

**YOUR GRADE:**

20% weekly problem sets

20% Quiz 1 (1 hour)

20% Quiz 2 (1 hour)

40% Final exam (3 hours)

### **Lecture plan:**

- L1-5: Background
- L6-13: Special relativistic kinematics
- L14-23: Special relativistic dynamics
- L24-25: Relativity of E&M
- L26-29: Cosmology
- L30-39: Curved spacetime & black holes

See the detailed syllabus for a lecture-by-lecture list of topics. The GR part will be limited; we'll mostly work out consequences of FRW and Schwarzschild metrics, and perhaps whet your appetite for a full GR course.

### **Reading:**

1. We'll start with Resnick (the whole book). On the side, browse French and (highly recommended!) Einstein. N.B. BOOK LATE!
2. For dynamics and E&M, we do French (5-8).
3. For cosmology, there will be handouts.
4. For black holes, we'll do Taylor & Wheeler.

**Consider printing the lecture notes beforehand and writing notes on the during class.**

# The History of Physics in 10 minutes

1. Ancient civilizations (what limited them?)
2. Newtonian Mechanics (1600's) (8.01, 8.06)
3. Electromagnetism (1800-1875) (8.02, 8.03, 8.07)
4. Stat Mech & Thermo (1850-1900)
5. Relativity (Einstein 1905, 1916) (8.033, 8.20, 8.224)
6. Quantum Mechanics (1900-1926)
7. 1900's breakthroughs
8. What's left for you to do?

## Newtonian Mechanics (1600's)

Tycho Brahe	1546-1601	Danish
Johannes Kepler	1571-1630	German
Sir Isaac Newton	1642-1727	English

*Principia* published 1678

## Era of gravitational astronomy (1700's)

Leonhard Euler	1707-1783	Swiss
Alexis Clairaut	1713-1765	French
J. D'Alembert	1717-1783	French
Joseph Lagrange	1736-1813	French
Pierre Laplace	1749-1827	French

## Era of electricity & magnetism (1800-1875)

Carl F. Gauss	1777-1855	German
André Ampere	1775-1836	French
Michael Faraday	1791-1867	English
Georg Ohm	1787-1854	German
James C. Maxwell	1831-1879	Scottish

## Statistical Mechanics & Thermodynamics (1850-1900)

Clausius	1822-1888
Joule	1818-1889
Kelvin	1824-1907
Helmholtz	1821-1871
Maxwell	1831-1879
Boltzmann	1844-1906
Planck	1858-1947

## Relativity

Albert Einstein 1899-1955

(Lorentz, Riemann, Schwarzschild, Kerr, FRWL, Wheeler, Kruskal, Hawking, ...)

## Quantum Mechanics

Max Planck 1858-1947

Niels Bohr 1885-1962

Louie de Broigle Max Born 1882-1970

Werner Heisenberg 1901-1976

Erwin Schödinger 1887-1961

Wolfgang Pauli 1900-1958

John von Neumann

Paul Dirac

Hugh Everett, III 1930-1982

Hans-Dieter Zeh

# Einstein

In 1905, at age 26, while working in a patent office with almost no contact to academia, he wrote three monumental papers:

- Photoelectric effect
- Brownian Motion
- Special theory of relativity

Completed general relativity in 1916. Learn from his approach!



# Hints of relativity in the physics you already know

## Special relativity

Maxwell's equations in vacuum imply

$$\nabla^2 \mathbf{E} - \frac{1}{c^2} \ddot{\mathbf{E}} = 0.$$

- This implies (as you learned in 8.02) waves traveling at speed  $c$  at *any* frequency (big shock then — now observed over range  $10^3 - 10^{27}$  Hz).
- But speed  $c$  relative to what? No reference to any particular frame.
- Consider flashlight on train. Either the equation is lacking something or something else is wrong. We'll explore what the big deal is.

# General relativity

(Please slow me down!)

Combining

$$F = ma$$

with

$$F = \frac{GmM}{r^2}$$



shows that the gravitational acceleration

$$a = \frac{GM}{r^2}$$

is mass-independent as long as

“inertial mass” = “gravitational mass”.

Is it?

- Galileo’s Pisa experiment showed it with low precision.
- Eötvös (1890) and later others showed with high precision that  $a$  independent of both mass and composition (density, atomic element, matter/antimatter, etc). Coincidence?
- In GR, Einstein explained it as gravity being a purely geometric effect. Follows from his equivalence principle.

# Key lessons of the course

## Special relativity

- Space and time are inextricably merged as 4D spacetime.
- Fast moving clocks appear slower, shorter and heavier.
- $E = mc^2$

## General relativity

- Spacetime is not static but dynamic, globally expanding and locally curving and contracting to form black holes *etc.*
- Matter curves spacetime
- Things moving straight through curved spacetime appear deflected (gravity)

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← (your theory here)

- Think for yourself.
- Question authority.
- Don't dismiss ideas just because they sound weird.

# *Expect* physics to feel weird!

- Relativistic driving movie
- Black hole movie

Next time:  
symmetry  
in physics