

Quasi-equilibrium Theory of Small Perturbations to Radiative- Convective Equilibrium States

- See *Quasi-Equilibrium Dynamics of the Tropical Atmosphere* paper on course web site
- Free troposphere assumed to have moist adiabatic lapse rate (s^* does not vary with height)
- Boundary layer quasi-equilibrium applies

Basis of statistical equilibrium physics

- Dates to Arakawa and Schubert (1974)
- Analogy to continuum hypothesis:
Perturbations must have space scales \gg intercloud spacing
- TKE consumption by convection \sim CAPE generation by large scale
- Numerical models on the verge of simulating clouds + large-scale waves
- We further assume convective criticality

Implications of the moist adiabatic lapse rate for the structure of tropical disturbances

- Approximate moist adiabatic condition as that of constant saturation entropy:

$$s^* = c_p \ln\left(\frac{T}{T_0}\right) - R_d \ln\left(\frac{p}{p_0}\right) + \frac{L_v q^*}{T}$$

- Assume hydrostatic perturbations:

$$\frac{\partial \phi'}{\partial p} = -\alpha'$$

- Maxwell's relation:

$$\alpha' = \left(\frac{\partial \alpha}{\partial s^*} \right)_p s^{*'} = \left(\frac{\partial T}{\partial p} \right)_{s^*} s^{*'}$$

- Integrate:

$$\phi' = \phi_b'(x, y, t) + \left(\bar{T}(x, y, t) - T \right) s^{*'}$$

Only barotropic and first baroclinic mode survive

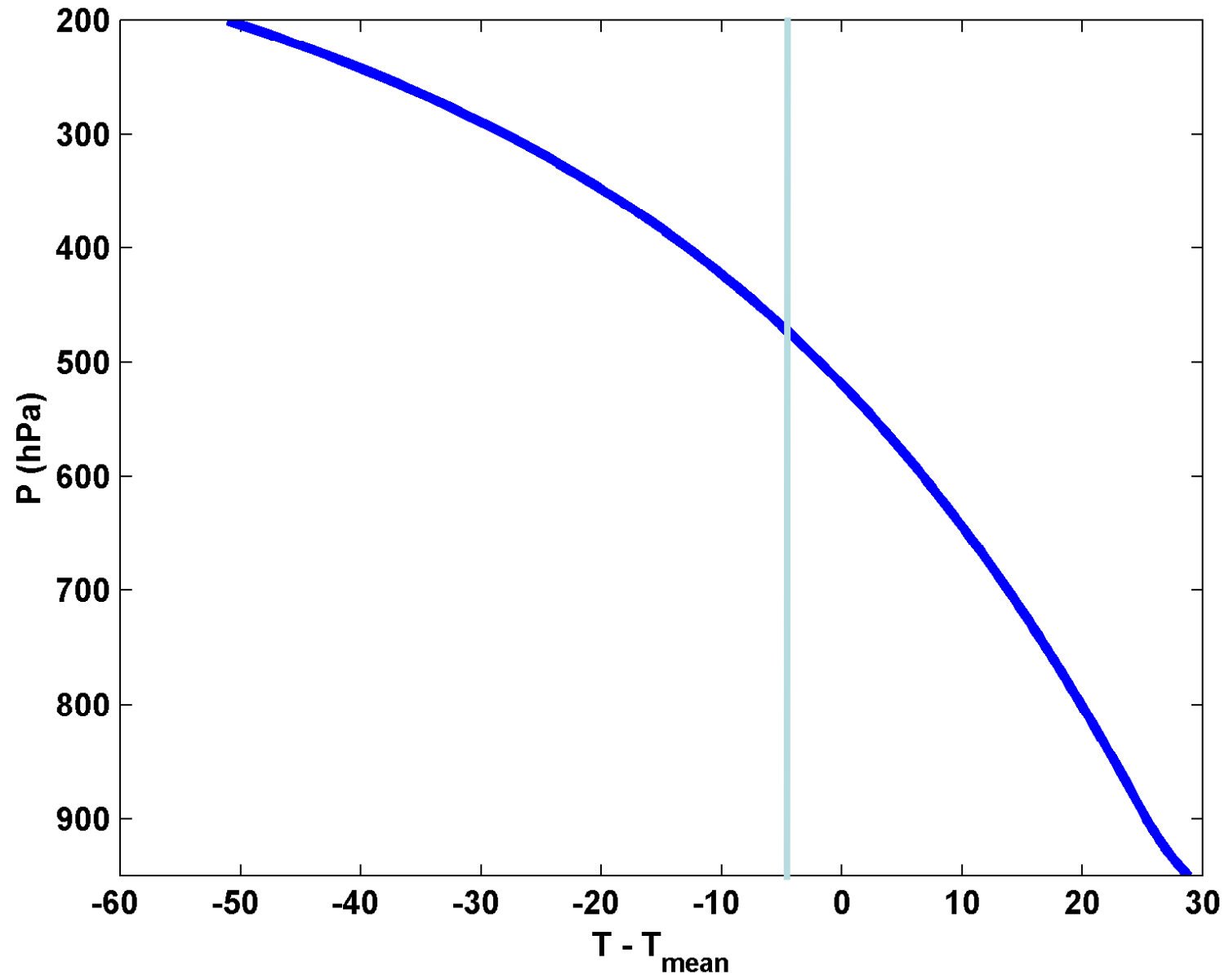
This implies, through the linearized momentum equations, e.g.

$$\frac{\partial u}{\partial t} = -\frac{\partial \phi}{\partial x} + fv$$

that the horizontal velocities may be partitioned similarly:

$$u = u_b(x, y, t) + \left(\bar{T}(x, y, t) - T \right) u^*(x, y, t);$$

$$v = v_b(x, y, t) + \left(\bar{T}(x, y, t) - T \right) v^*(x, y, t).$$



Implications for vertical structure of vertical velocity

$$\frac{\partial \omega}{\partial p} = - \left(\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} \right)$$

Integrate:

$$\omega = (p_0 - p) \left(\frac{\partial u_b}{\partial x} + \frac{\partial v_b}{\partial y} \right) - \left((p_0 - p) \bar{T} - \int_p^{p_0} T dp \right) \left(\frac{\partial u^*}{\partial x} + \frac{\partial v^*}{\partial y} \right).$$

At tropopause:

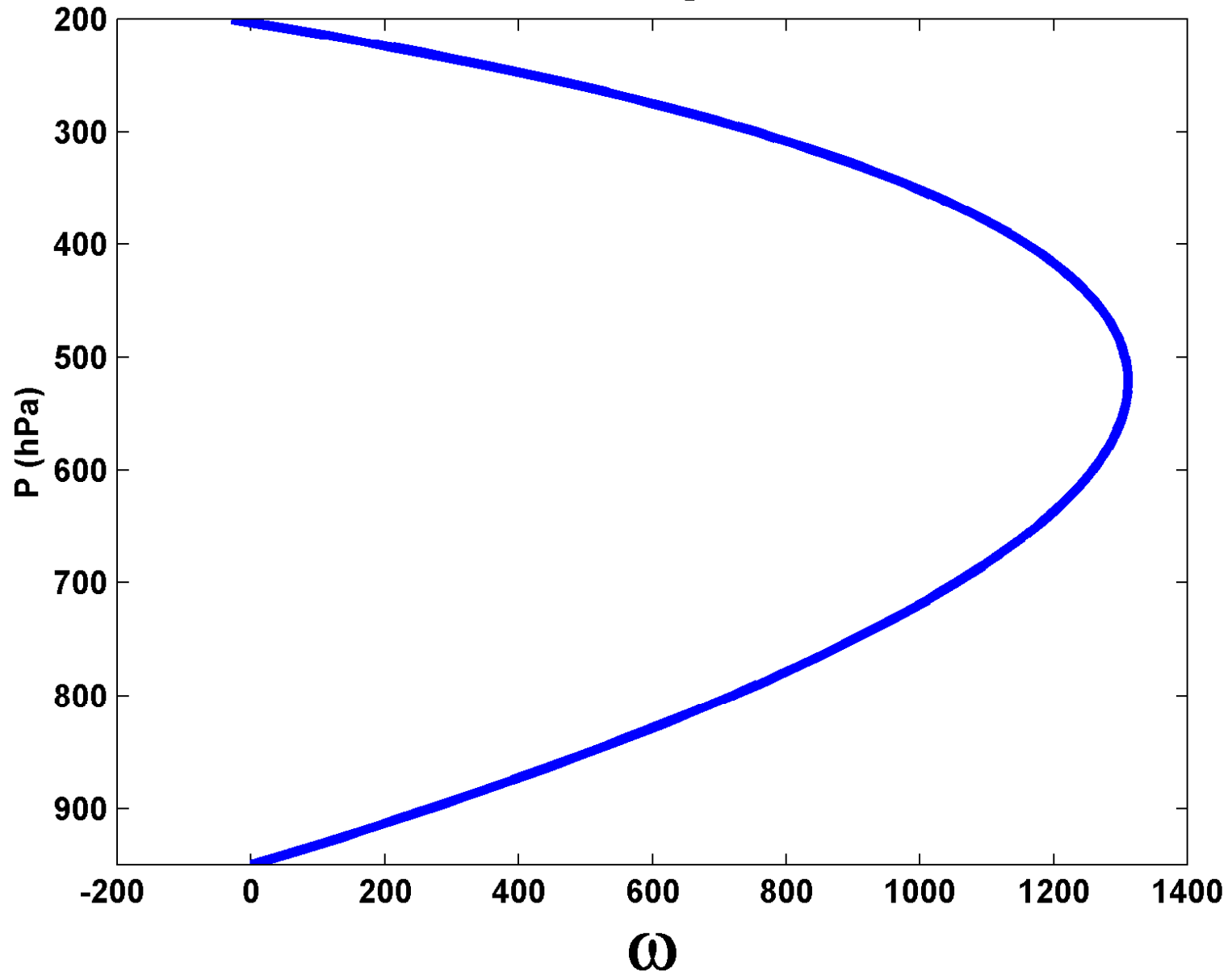
$$\omega_t = (p_0 - p_t) \left(\frac{\partial u_b}{\partial x} + \frac{\partial v_b}{\partial y} \right)$$

This implies that if a rigid lid is imposed at the tropopause, the divergence of the barotropic velocities must vanish and the barotropic components therefore satisfy the barotropic vorticity equation:

$$\frac{\partial \eta_b}{\partial t} = -\mathbf{V} \cdot \nabla \eta_b,$$

$$\eta_b \equiv \hat{k} \cdot \nabla \times \mathbf{V}_b + 2\Omega \sin \theta$$

Omega



Feedback of Air Motion on (virtual) Temperature

- Convection cannot change vertically integrated enthalpy, $k = c_p T + L_v q$
- Then neglecting surface fluxes, radiation, and horizontal advection,

$$\frac{\partial}{\partial t} \int k dp = - \int \omega \frac{\partial h}{\partial p} dp,$$

- Neelin and Held (1987): This function is negative for upward motion

- Upward motion is associated with column moistening:

$$\int c_p \frac{\partial T}{\partial t} dp = \frac{\partial}{\partial t} \int k dp - \int L_v \frac{\partial q}{\partial t} dp$$

—————> **Ascent leads to cooling**

- Yano and Emanuel, 1991:

$$N_{eff}^2 = (1 - \varepsilon_p) N^2$$

Prediction: Inviscid, small amplitude perturbations under rigid lid: Shallow water solutions with reduced equivalent depth

Quasi-Linear β Plane System , Neglecting Barotropic Mode

$$\frac{\partial u}{\partial t} = (T_s - \bar{T}) \frac{\partial s^*}{\partial x} + \beta yv - ru$$

$$\frac{\partial v}{\partial t} = (T_s - \bar{T}) \frac{\partial s^*}{\partial y} - \beta yu - rv$$

$$\frac{\partial s^*}{\partial t} = \frac{\Gamma_d}{\Gamma_m} \left(\dot{Q}_{rad} + \frac{\partial s_d}{\partial z} (\varepsilon_p M - w) \right)$$

$$h \frac{\partial s_b}{\partial t} = C_k |\mathbf{V}| (s_0^* - s_b) - (M - w)(s_b - s_m)$$

$$\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} + \frac{w}{H} = 0$$

Quasi-Equilibrium Assumption:

$$\frac{\partial s_b}{\partial t} = \frac{\partial s^*}{\partial t}$$

Gives closure for convective mass flux, M

System closed except for specification of

$$\dot{Q}_{rad}, s_0^*, s_m, \varepsilon_p$$

Additional Approximations:

- **Boundary Layer QE** (Raymond, 1995): Neglect $h \frac{\partial s_b}{\partial t}$, gives simpler expression for M

$$M = w + C_k |\mathbf{V}| \frac{s_0^* - s_b}{s_b - s_m}$$

- **Weak Temperature Approximation** (Sobel and Bretherton, 2000): Neglect $\frac{\partial s^*}{\partial t}$ (Over-determined system, ignore momentum equation for irrotational flow)

Important Feedbacks:

- **Wind-Induced Surface Heat Exchange (WISHE)** Coupling of surface enthalpy flux to wind perturbations (Neelin et al. 1987, Emanuel, 1987)
- **Moisture-Convection Feedback:**
Dependence of s_m on M and/or ε_p on $s^* - s_m$
- **Cloud-Radiation Feedback:**
Dependence of \dot{Q}_{rad} on M or $s^* - s_m$

- **Ocean-Atmosphere Feedback** (e.g. ENSO): Feedback between perturbation surface wind and ocean surface temperature, as represented by s_0^*

Simple Example:

- Ignore perturbations of \dot{Q}_{rad}
- Ignore fluctuations of ε_p
- Make boundary layer QE approximation
- Fully linearize surface fluxes:

$$|\overline{\mathbf{V}}| = \sqrt{\overline{U}^2 + u^{*2}}$$

$$|\mathbf{V}'| = \frac{\overline{U}u'}{|\overline{\mathbf{V}}|}$$

Introduce scalings:

First define a meridional scale, L_y :

$$L_y^4 = \frac{\Gamma_d}{\Gamma_m} (T_s - \bar{T}) H \frac{\partial s_d}{\partial z} \frac{1 - \varepsilon_p}{\beta^2}$$

Then let

$$x \rightarrow a x$$

$$y \rightarrow L_y y$$

$$t \rightarrow \frac{a}{\beta L_y^2} t$$

$$u \rightarrow \frac{a C_k \overline{|\mathbf{V}|}}{H} u$$

$$v \rightarrow \frac{L_y C_k \overline{|\mathbf{V}|}}{H} v$$

$$s^* \rightarrow \frac{a C_k \overline{|\mathbf{V}|} \beta L_y^2}{H (T_s - \bar{T})} s^*$$

Separate scalings for ocean temperature and lower tropospheric entropy:

$$s_o^* \rightarrow \frac{1 - \epsilon_p}{\epsilon_p} \left(\overline{s_b - s_m} \right) s_0$$

$$s_m \rightarrow \frac{1 - \epsilon_p}{\epsilon_p} \frac{\left(\overline{s_b - s_m} \right)^2}{\left(\overline{s_o^* - s^*} \right)} s_m$$

Nondimensional parameters:

$$\alpha \equiv \frac{1 - \varepsilon_p}{\varepsilon_p} \frac{aC_k}{H} \frac{\bar{U}}{|\mathbf{V}|} \frac{\left(\overline{s_0^* - s^*}\right)}{\left(\overline{s^* - s_m}\right)} \quad (\text{WISHE})$$

$$\mathcal{R} \equiv \frac{ra}{\beta L_y^2} \quad (\text{Rayleigh friction})$$

$$\chi \equiv \frac{1 - \varepsilon_p}{\varepsilon_p} \frac{aC_k |\mathbf{V}| \beta L_y^2 \left(\overline{s_0^* - s^*}\right)}{H (T_s - \bar{T}) \left(\overline{s^* - s_m}\right)^2} \quad (\text{surface damping})$$

$$\delta = \left(\frac{a}{L_y}\right)^2 \quad (\text{zonal geostrophy})$$

Nondimensional Equations:

$$\frac{\partial u}{\partial t} = \frac{\partial s}{\partial x} + yv - \mathcal{R}u$$

$$\frac{\partial v}{\partial t} = \delta \left(\frac{\partial s}{\partial y} - yu - \mathcal{R}v \right)$$

$$\frac{\partial s}{\partial t} = \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} + \alpha u + s_0 + s_m - \chi s$$

Steady System with $\mathcal{R} = s_m = 0$:

$$\frac{\partial s}{\partial x} + \alpha y \frac{\partial s}{\partial y} - \chi y^2 s = -y^2 s_0$$

Similar to Gill Model, but forcing is directly in terms of SST (s_0), not latent heating

For SST of the form

$$s_0 = \mathbf{RE} \left[G(y) e^{ikx} \right]$$

there are solutions of the form

$$s = \mathbf{RE} \left[J(y) e^{ikx} \right],$$

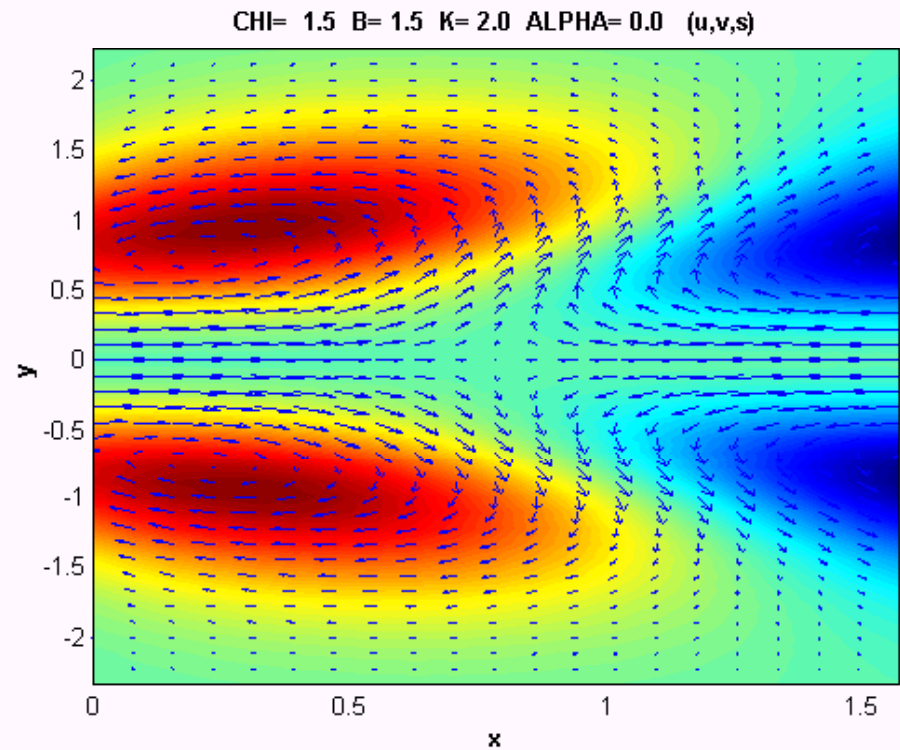
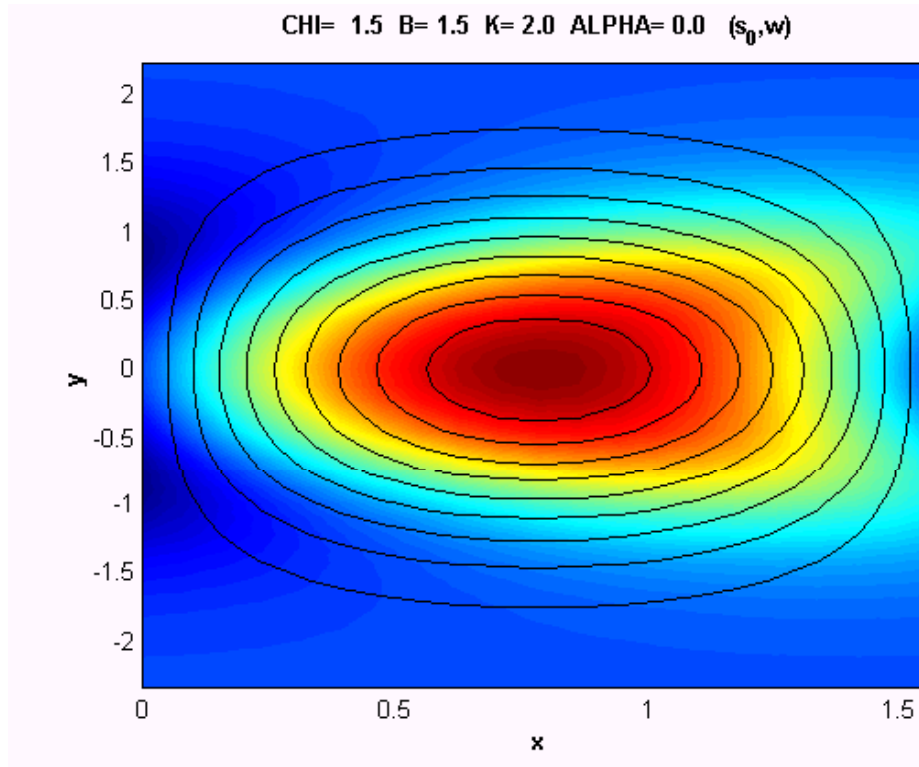
where

$$J(y) = y^{-ik/\alpha} e^{\chi y^2/2\alpha} \int_0^y G u^{1+ik/\alpha} e^{-\chi u^2/2\alpha} du$$

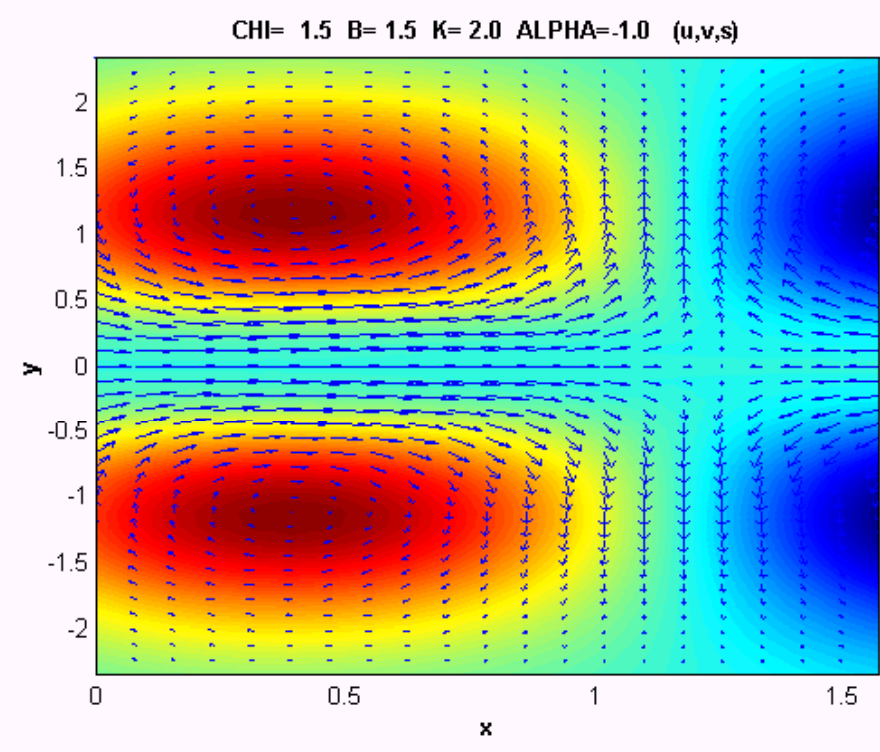
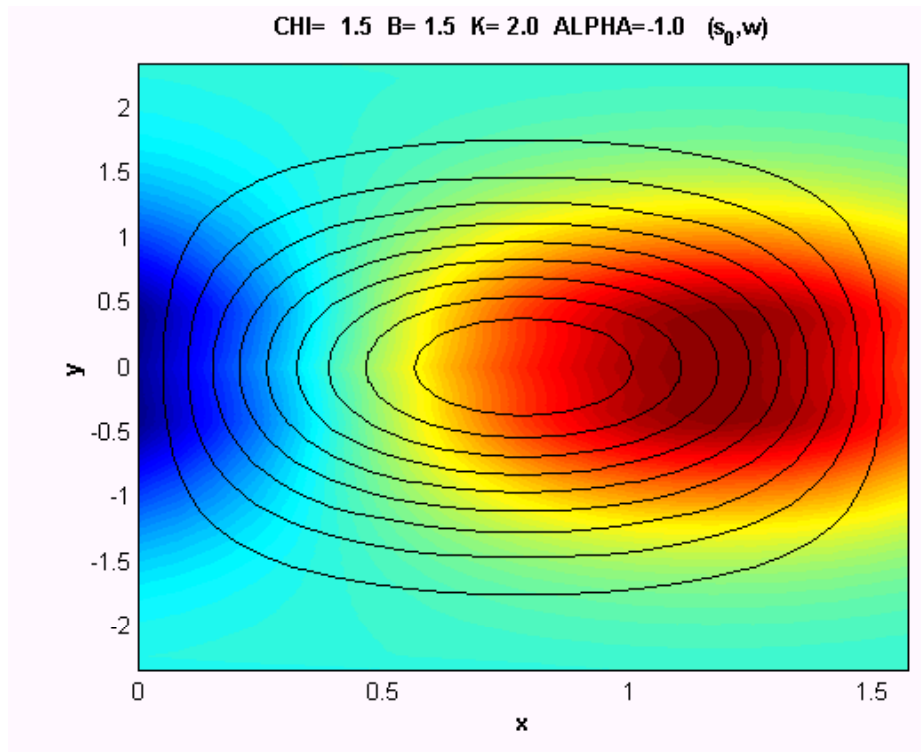
Example:

$$G = e^{-by^2}$$

$$\alpha = 0, \quad k = 2, \quad b = 1.5, \quad \chi = 1.5$$



$$\alpha = -1, \quad k = 2, \quad b = 1.5, \quad \chi = 1.5$$



Basic linear wave dynamics on the equatorial β plane

Omit damping and WISHE terms from linear nondimensional equations:

$$\frac{\partial u}{\partial t} = \frac{\partial s}{\partial x} + yv$$

$$\frac{\partial v}{\partial t} = \delta \left(\frac{\partial s}{\partial y} - yu \right)$$

$$\frac{\partial s}{\partial t} = \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y}$$

Fully equivalent to the shallow water equations on a β plane

Eliminate s and u in favor of v :

$$\frac{\partial}{\partial t} \left\{ \frac{\partial^2 v}{\partial t^2} - \frac{\partial^2 v}{\partial x^2} - \delta \frac{\partial^2 v}{\partial y^2} + \delta y^2 v \right\} - \delta \frac{\partial v}{\partial x} = 0$$

Let $v = V(y)e^{ikx - i\omega t}$

$$\rightarrow \frac{d^2 V}{dy^2} + \left(\frac{\omega^2 - k^2}{\delta} - \frac{k}{\omega} - y^2 \right) V = 0$$

Boundary conditions: V well behaved at $y \rightarrow \pm\infty$

Solution in terms of discrete parabolic cylinder functions D_n :

$$v = D_n(y),$$

$$\text{where } D_n = e^{-2y^2} [1, 2y, 4y^2 - 2, \dots]$$

provided ω satisfies the dispersion relation

$$\boxed{\frac{\omega^2 - k^2}{\delta} - \frac{k}{\omega} = 2n + 1}$$

There is, in addition, another mode satisfying $v=0$ everywhere. From first and third linear equations:

$$\frac{\partial^2 u}{\partial t^2} - \frac{\partial^2 u}{\partial x^2} = 0.$$

Satisfied by $u = F(x - t).$

Eastward-propagating, nondispersive
equatorially trapped Kelvin wave

Note that this happens to satisfy derived
dispersion relation when $n = -1$.

There are three roots of the general dispersion relation:

$$n = 0: \quad \frac{\omega^2 - k^2}{\delta} - \frac{k}{\omega} - 1 = 0$$

$$\text{Factor:} \quad \left(\frac{\omega - k}{\delta} - \frac{1}{\omega} \right) (\omega + k) = 0$$

$\omega = -k$ root not allowed (*does not satisfy BCs*)

$$\omega = \frac{1}{2} \left(k \pm \sqrt{k^2 + 4\delta} \right)$$

Mixed Rossby-Gravity Waves (MRG)

For $n \geq 1$, two well defined limits:

1. $|\omega| \ll |k|$:

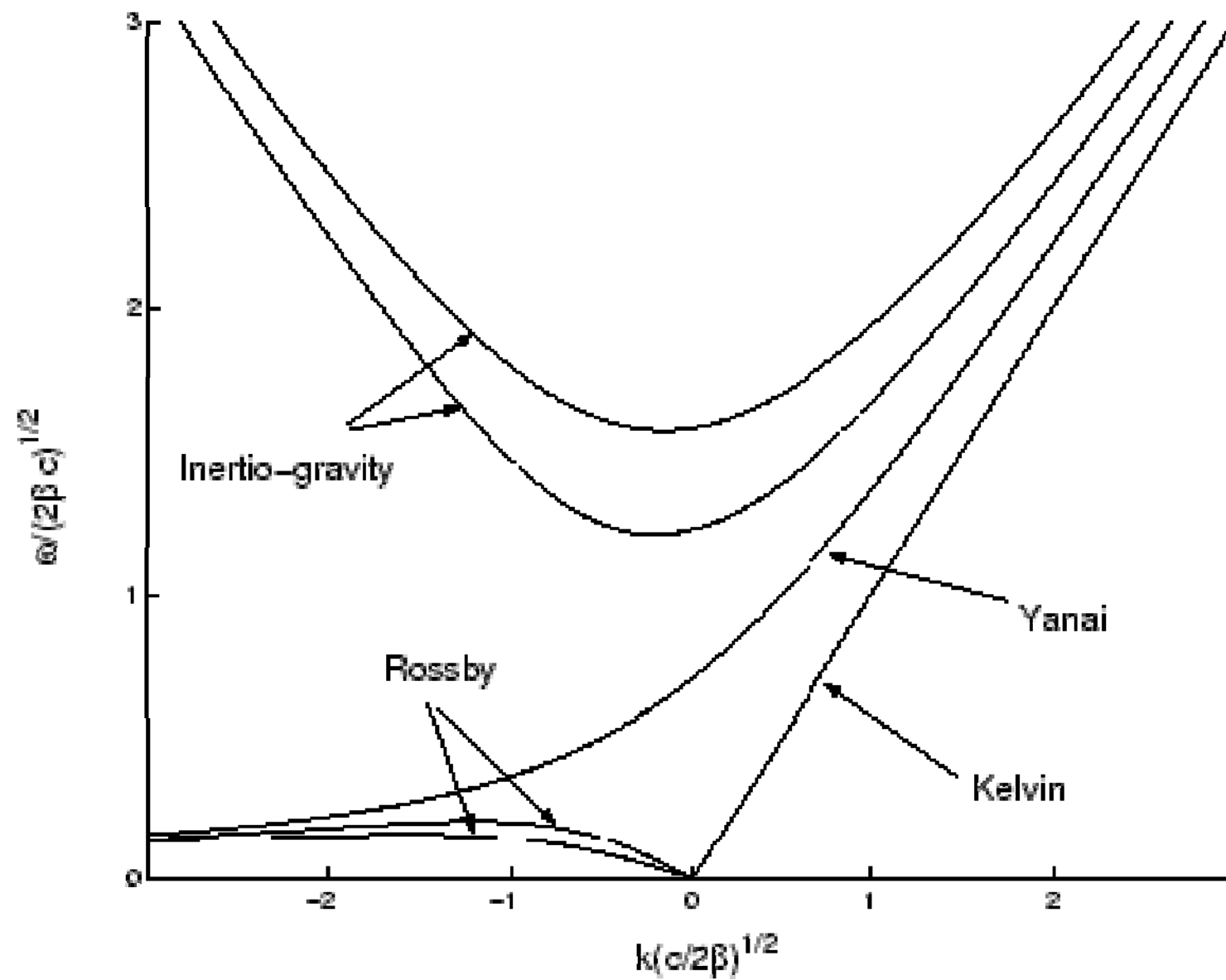
$$\omega \cong -\frac{k}{2n+1 + k^2/\delta}$$

Planetary
Rossby waves

2. $|\omega| \gg |k|$:

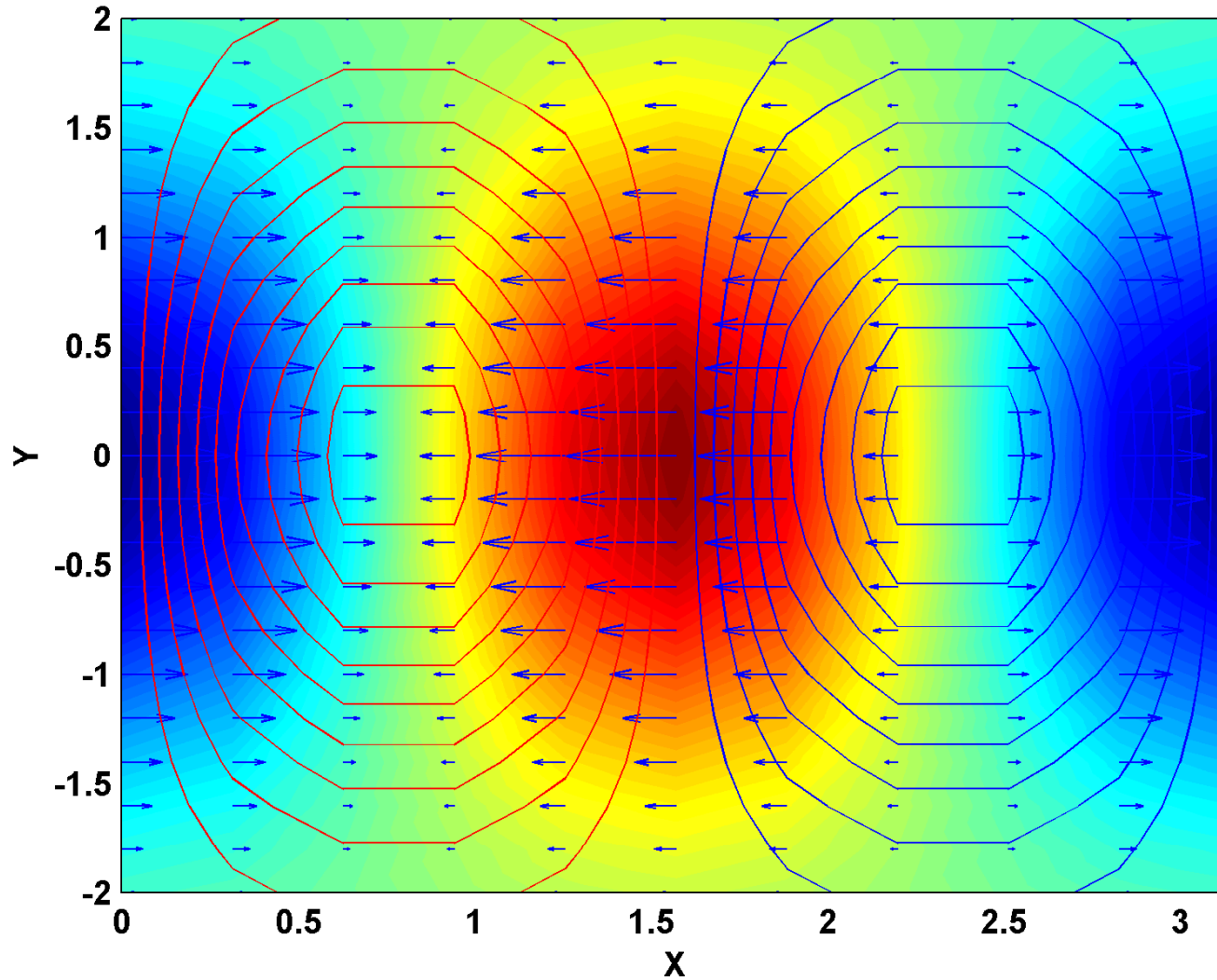
$$\omega^2 \cong k^2 + \delta(2n+1)$$

Inertia-gravity
waves



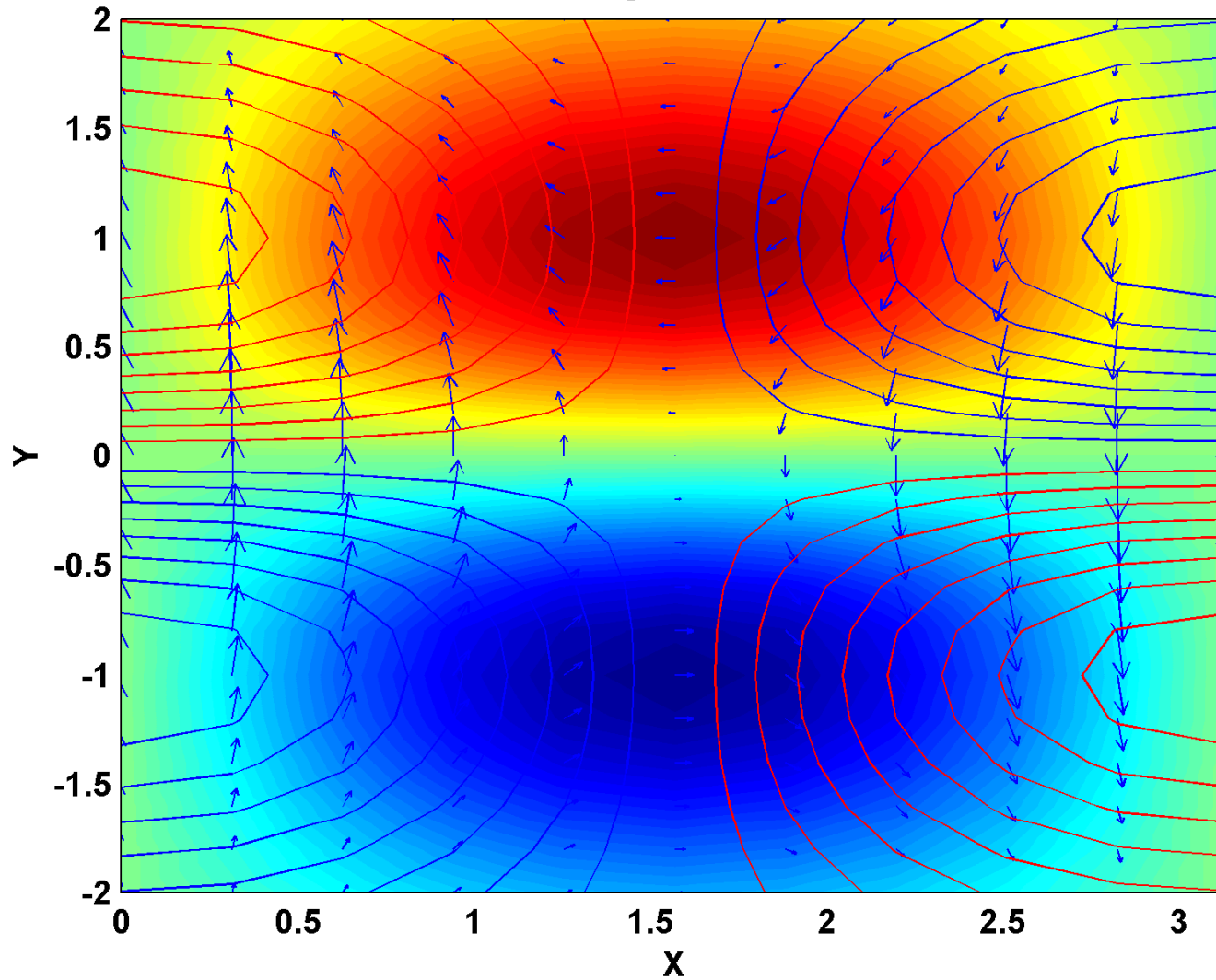
Kelvin wave

s, u, v, and w, kel, n= -1 k= 2 delta= 10



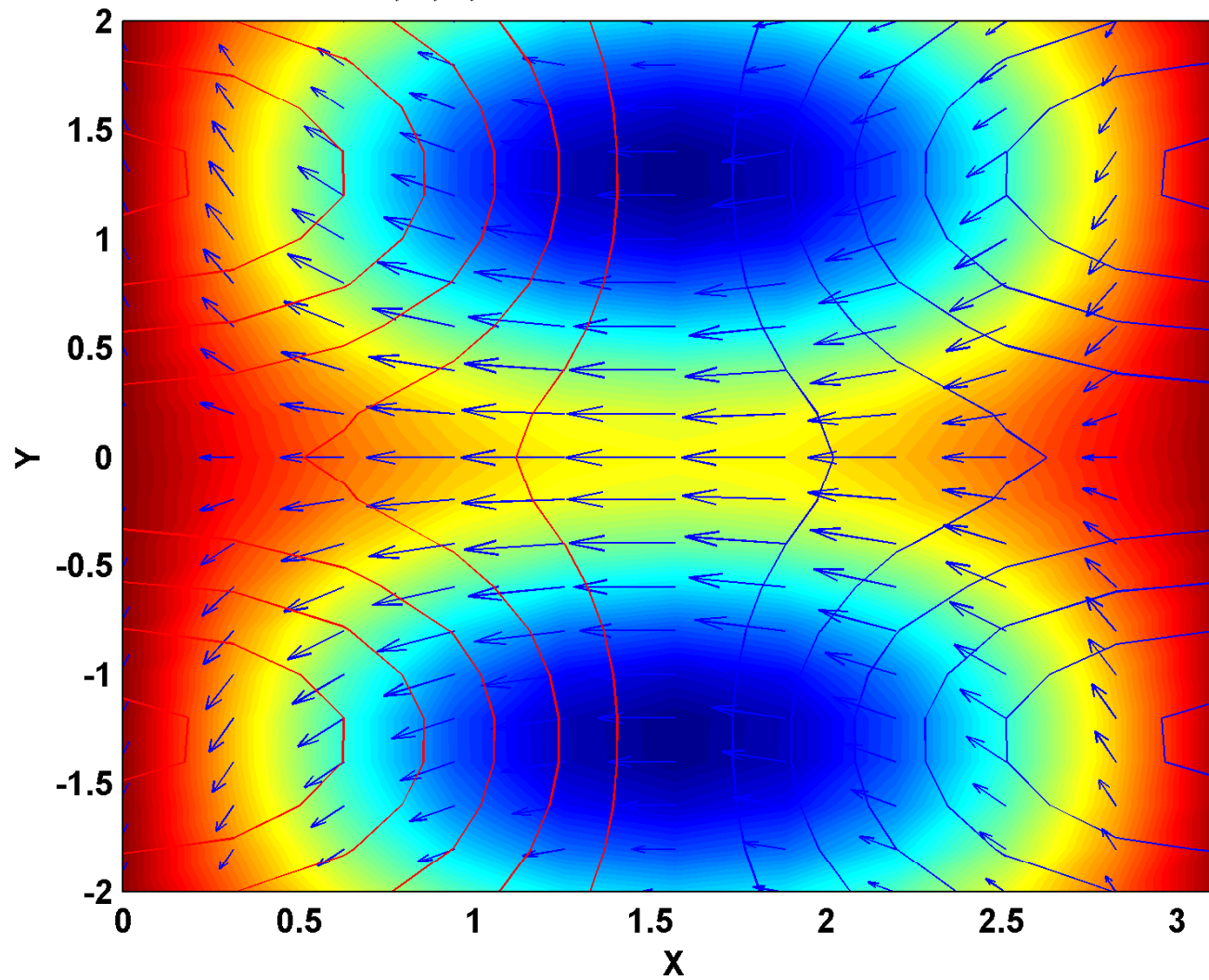
Mixed Rossby-Gravity

s, u, v, and w, mrg, n= 0 k= -1 delta= 10



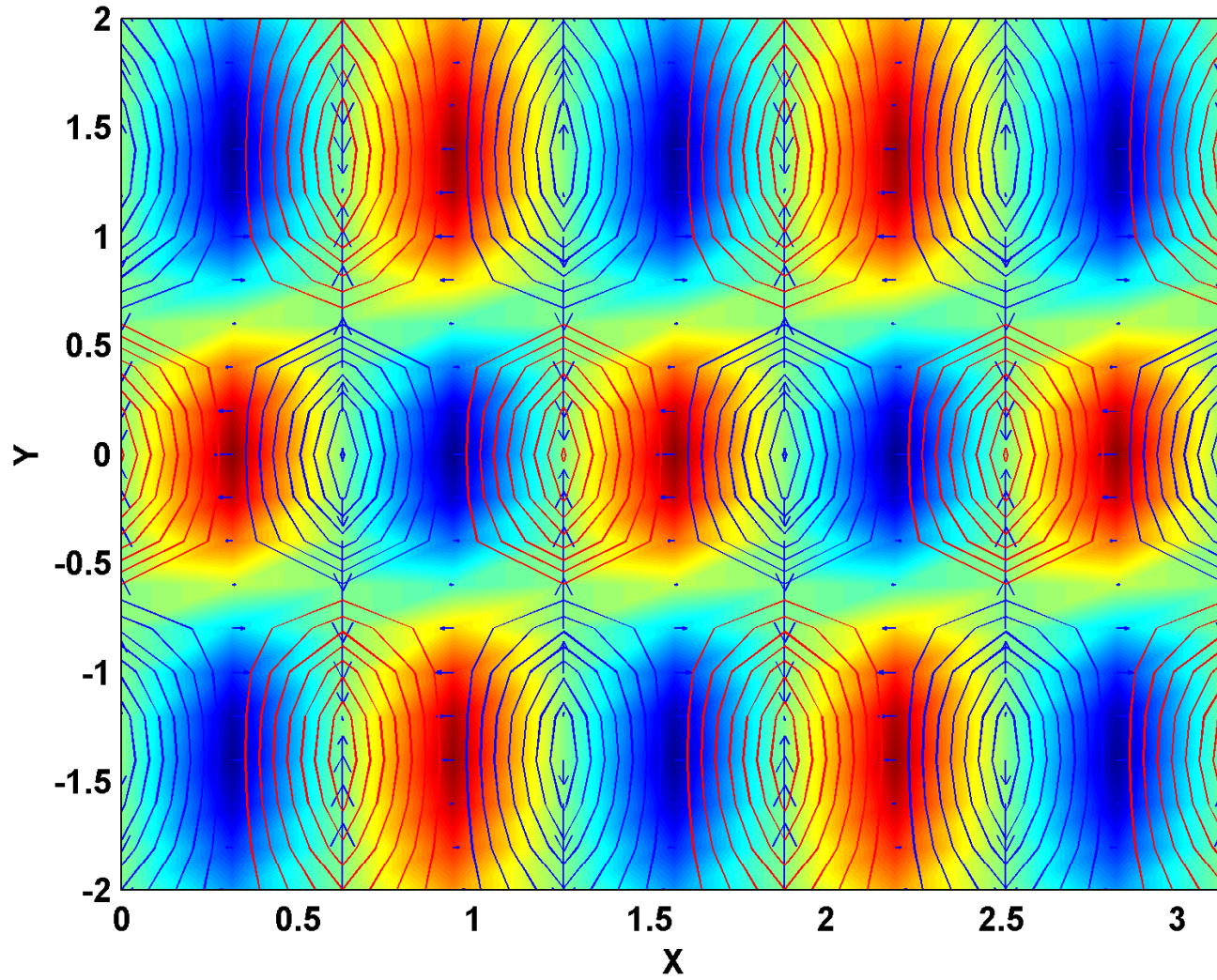
Rossby

s, u, v, and w ros n= 1 k= 1 delta= 10



Inertia-gravity

s, u, v, and w, ing, n= 3 k= 5 delta= 10



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